

# Theory Overview of Chiral Magnetic Effect in Heavy-Ion Collisions

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(Acknowledgment: CME Task Force Meeting)

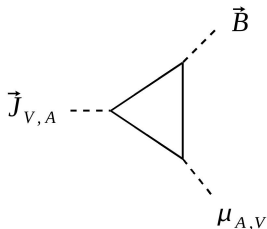
# Plan of the talk

- A question: **Is the Chiral Magnetic Wave (CMW) something different from the Chiral Magnetic Effect (CME)?**
- **Issues of CME in heavy-ion collisions**
  - Issues of CME itself
  - Issues related to heavy-ion collision
  - Issues related to other backgrounds of the proposed observables
- **What do we need to do?**

# Chiral Magnetic Effect (CME) and Chiral Separation Effect (CSE)

(Fukushima-Kharzeev-Warringa, Son-Zhitnitsky, Vilenkin)

$$\vec{J}_V = \frac{eN_c}{2\pi^2} \mu_A \vec{B}, \quad \vec{J}_A = \frac{eN_c}{2\pi^2} \mu_V \vec{B}$$



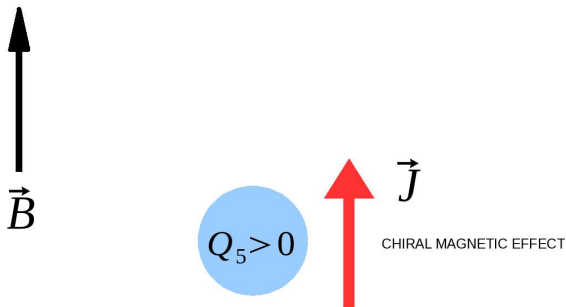
**Note the  $\langle AVV \rangle$  structure**

**Q: Is the Chiral Magnetic Wave (CMW) something different from the Chiral Magnetic Effect (CME)?**

**A:  $\text{CME} + \text{CSE} + \text{Hydrodynamics} = \text{CMW}$**

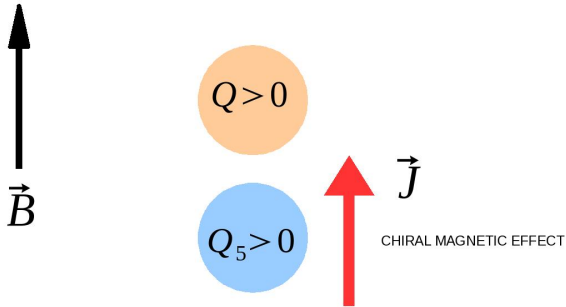
**CMW is the inclusive universal language of CME/CSE in hydrodynamics**

# Why do we have waves ?



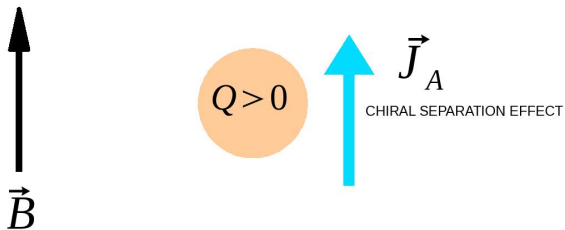
$$\vec{J}_V = \frac{N_c e \vec{B}}{2\pi^2} \mu_A \quad , \quad \vec{J}_A = \frac{N_c e \vec{B}}{2\pi^2} \mu_V$$

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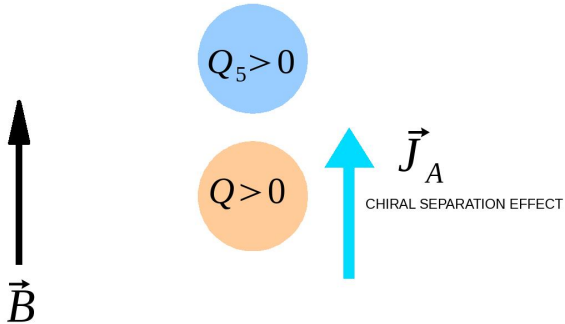
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# CMW: Sound Waves of Chiral Charges (Kharzeev-HUY)

Add and subtract CME and CSE to go to  
**Left/Right-handed chiralities**

$$\vec{J}_V = \frac{eN_c}{2\pi^2} \mu_A \vec{B}, \quad \vec{J}_A = \frac{eN_c}{2\pi^2} \mu_V \vec{B}$$

$$J_{R/L} \equiv \frac{1}{2}(J_V \pm J_A)$$

Then, we have a “diagonalization” of the CME/CSE

$$\vec{J}_R = \frac{eN_c}{4\pi^2} \mu_R \vec{B}, \quad \vec{J}_L = -\frac{eN_c}{4\pi^2} \mu_L \vec{B}, \quad \mu_{L/R} \approx \frac{1}{\alpha} n_{L/R}$$

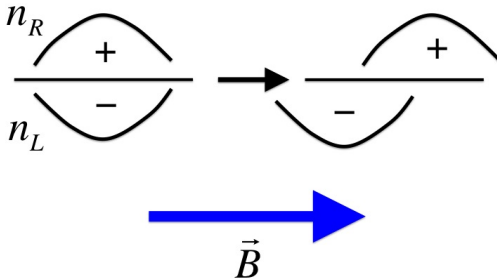
Hydro charge conservation  $\partial_\mu J_{L/R}^\mu = 0$  gives

$$\left( \partial_t + \vec{v}_\chi \cdot \vec{\nabla} \right) n_R = 0, \quad \left( \partial_t - \vec{v}_\chi \cdot \vec{\nabla} \right) n_L = 0, \quad \vec{v}_\chi = \frac{eN_c}{4\pi^2 \alpha} \vec{B}$$

**Two Independent Uni-directional Propagating Waves !**

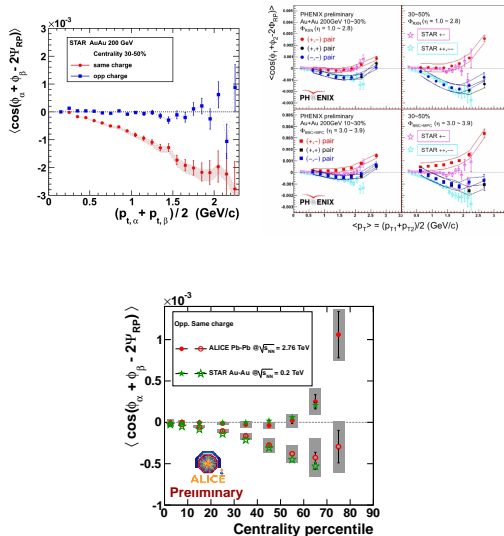
# Development of a **charge dipole** from initial axial charge

$$n_A > 0, \quad n_V = 0 \longrightarrow n_R = \frac{1}{2}n_A, \quad n_L = -\frac{1}{2}n_A = -n_R$$



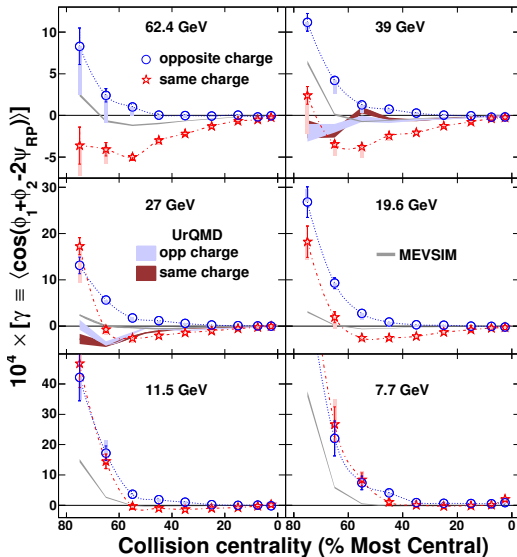
# Charge Dipole Probe in RHIC and LHC

$$\langle \cos(\phi_1 + \phi_2 - 2\Psi_{RP}) \rangle$$



# Beam Energy Scan

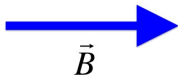
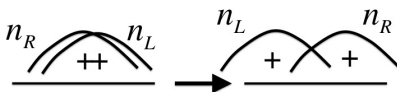
(Phys.Rev.Lett. 113 (2014) 052302)



# Development of a **charge quadrupole** from initial electric charge

(Burnier-Kharzeev-Liao-HUY, Gorba-Miransky-Shovkovy)

$$n_A = 0, \quad n_V > 0 \longrightarrow n_R = \frac{1}{2}n_V, \quad n_L = \frac{1}{2}n_V = n_R$$

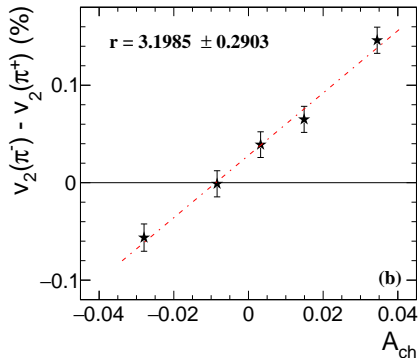
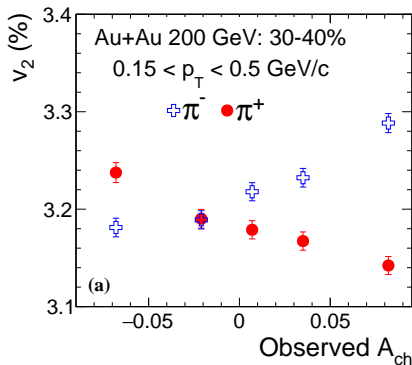


**Prediction:**

$$v_2(\pi^-) - v_2(\pi^+) = r Q_{init} \text{ with the slope } r > 0$$

# Linear dependency check

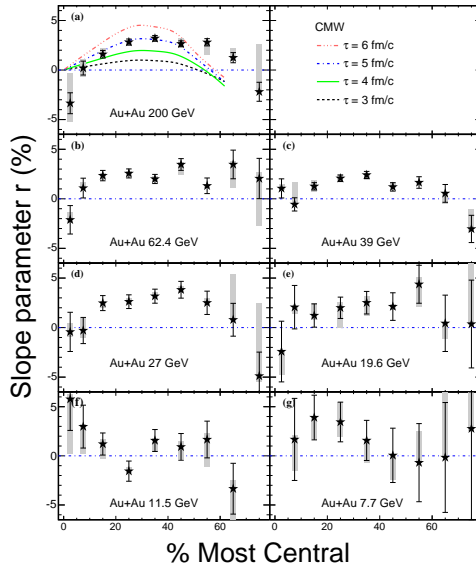
(Phys.Rev.Lett. 114 (2015) 25, 252302)



$\Delta v_2 = r A_{\pm} + \Delta v_2^0$ : Note the intercept  $\Delta v_2^0$  (Stephanov-HUY)

# Beam energy Scan

(Phys.Rev.Lett. 114 (2015) 25, 252302)



**A reasonable expectation is:**

**What should be happening is a mixture of  
dipole and quadrupole**

**CMW is the universal language for both  
phenomena**



## Four Commonly Raised Issues in CME

- **Q:** Axial charge is not conserved. Does CME make sense?
- **Q:** Electromagnetic fields are dynamical. Should't we solve the Maxwell equations with CME?
- **Q:** Magnetic field is rapidly changing in time. Does CME remain the same?
- **Q:** How much do we know about CME in the early pre-equilibrium phase of the QGP?

## Q: Axial charge is not conserved

Conservation law is modified to  $\partial_\mu J_A^\mu = -\frac{1}{\tau_R} n_A$  where the relaxation rate is given by

$$\frac{1}{\tau_R} \sim \alpha_s^5 \log(1/\alpha_s) T + \alpha_s m_q^2 / T$$

The CMW dispersion relation becomes  
(Stephanov-HUY-Yin)

$$\omega = -\frac{i}{2\tau_S} \pm \sqrt{-\frac{1}{4\tau_S^2} + v_x^2 k^2}$$

A transition from CMW ( $k \gg 1/\tau_R$ ) to pure diffusion  
( $k \ll 1/\tau_R$ )

$\tau_R$  is numerically larger than 10 fm with  $\alpha_s = 0.2$  and  
 $m_q = 10$  MeV

It is okay to neglect this in heavy-ion collisions

**Q: Electromagnetic field is dynamical**  
**Longitudinal Mode:**

The CMW dispersion becomes (Kharzeev-HUY)

$$\omega^2 = v_\chi^2 k^2 + m_B^2$$

where

$$v_\chi = \frac{eN_c}{2\pi^2\alpha} B, \quad m_B \equiv \frac{e^2 N_c}{2\pi^2 \sqrt{\alpha}} B$$

**CMW becomes “massive”**

Numerically,  $m_B^{-1} \approx 20$  fm with  $eB \sim m_\pi^2$  and  $T \approx 200$  MeV, the basic reason is the smallness of  $\alpha_{EM}$

**It is okay to neglect this in heavy-ion collisions**

## Q: **Electromagnetic field is dynamical**

### **Transverse Mode:**

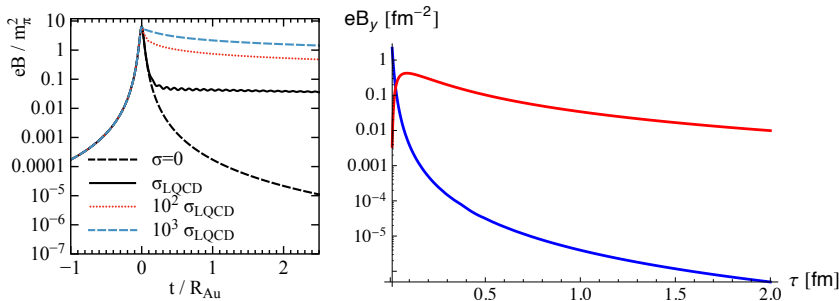
Transverse modes become helicity-polarized via the transition from the fermion helicity (axial charge) into the magnetic helicity. For low enough  $k$ , the resulting chiral magneto-hydrodynamics features inverse cascades with turbulence (eg. a recent work by Hirono-Kharzeev-Yin and Yamamoto).

Due to the smallness of  $\alpha_{EM}$  and  $\mu_A/T$ , the space-time scale  $k \sim \alpha_{EM}\mu_A$  is numerically much larger than 10 fm

**It is okay to neglect this in heavy-ion collisions,  
unless  $\mu_A/T \sim 1/\alpha_{EM} \sim 100$**

# Q: Magnetic field is rapidly changing

## Magnetic field in QGP medium



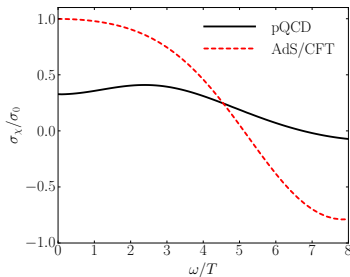
(Tuchin, McLerran-Skokov, Gursoy-Kharzeev-Rajagopal)

**Faraday Effect with equilibrium QGP conductivity, which can be questioned**

# CME in the time-changing magnetic field

In frequency space,

$$\vec{J}(\omega) = \sigma_\chi(\omega) \vec{B}(\omega)$$



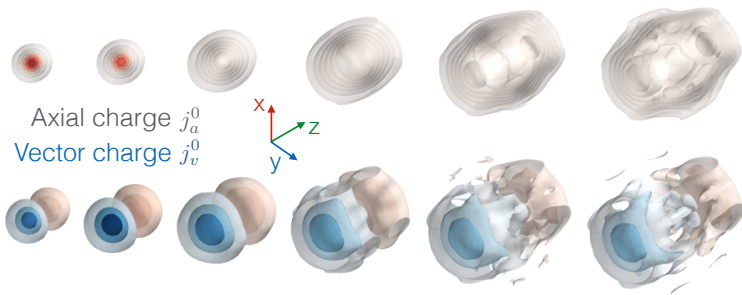
(Kharzeev-Warringa, HUY, Jimenez-HUY)

$$\sigma_\chi^{\text{pQCD}}(\omega) \approx \sigma_0 \left( 1 - \frac{2}{3} \frac{\omega}{\omega + i\tau_R^{-1}} \right), \quad \tau_R^{-1} \approx \frac{1}{36} g^4 \log(1/g) T$$

**Q: CME/CMW in Pre-Equilibrium**

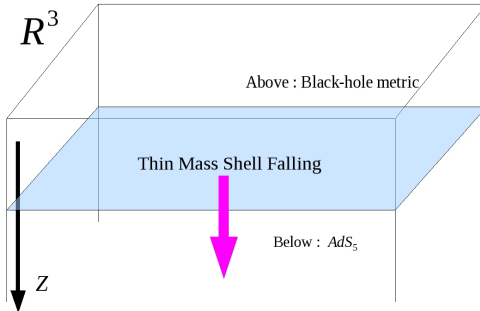
**A: We don't know well enough**

**Recent real-time lattice simulation:**  
(Mueller-Schlichting-Sharma)



# AdS/CFT: (Lin-HUY)

(Initial) Thermalization (2) : Falling Mass Shell



(Lin-Shuryak)



# Dispersion Relation

$$\omega = \sqrt{f(z_s)}k + \Delta\omega$$

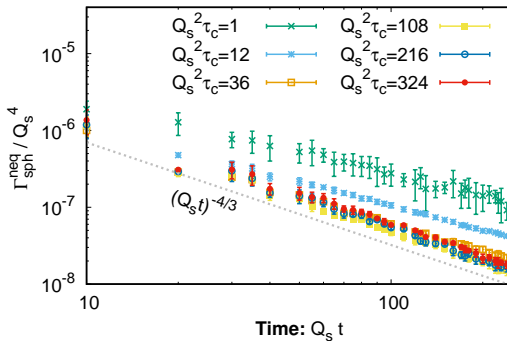
**where  $z_s$  is the holographic location of the mass shell, and  $f(z_s) \approx 1$  initially and  $f(z_s) \rightarrow 0$  when  $z_s \rightarrow z_H$  (horizon)**

**$\Delta\omega$  features characteristics of CMW, and CMW velocity is enhanced by  $\sqrt{f(z_s)}$**

$$v_{\chi}^{\text{PreEq}} = \sqrt{f(z_s)} + v_{\chi}^{\text{Eq}}, \quad v_{\chi}^{\text{Eq}} \equiv \frac{\partial \Delta\omega}{\partial k}$$

In addition: The amount of initial axial charge is difficult to compute

Recent Glasma simulation: (Mace-Schlichting-Venugopalan)



**Other background effects to the charge dipole and quadrupole observables exist, and are sizable**

**For the charge dipole observable:** Local charge conservation and transverse momentum conservation give additional correlations to the two particles (**Pratt-Schlichting, Bzdak-Koch-Liao**)

**For the charge quadrupole observable:** Iso-spin transports and viscous corrections give additional contribution to the slope parameter  $r$  (**Dunlop-Lisa-Sorensen, Hatta-Monnai-Xiao**)

## Q: What should be done next ?

- Devise a way to understand the pre-equilibrium axial dynamics and CME:
  - Interpolate between real-time AdS/CFT and real-time lattice simulations
  - Simulate Vlasov-Wong equation with Chiral Kinetic Theory (Son-Yamamoto, Stephanov-Yin) in the Glasma
- Run realistic anomalous hydro simulations including all possible background effects (the talk by Yuji Hirono):  
It is planned in the Beam Energy Scan Theory (BEST) topical collaboration

**Thank you very much for listening**